

# Physical Interpretation of PP-waves With Axial Torsion

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A spacetime with pp-metric and torsion

$$T := *A \tag{1}$$

where  $A$  is a real vector field defined by  $A = k(\varphi)l$ .

(*V. Pasic and E. Barakovic: "Torsion wave solutions in Yang-Mielke theory of gravity"*, Advances in High Energy Physics, accepted for publication. )

The torsion  $T$  is purely axial and the connection  $\Gamma$  is metric compatible.

The remarkable property: the curvature of a generalised pp-wave is a sum of the curvature of the underlying classical pp-space

$$-\frac{1}{2}(l \wedge \{\nabla\}) \otimes (l \wedge \{\nabla\})f \quad (2)$$

and the curvature

$$\frac{1}{4}k(\varphi)^2 \text{Re}((l \wedge m) \otimes (l \wedge \bar{m})) \mp \frac{1}{2}k'(\varphi) \text{Im}((l \wedge m) \otimes (l \wedge \bar{m})) \quad (3)$$

generated by a axial torsion wave traveling over the pp-space.

Ricci curvature is

$$\text{Ric} = \frac{1}{2}(f_{11} + f_{22} - k^2)(l \otimes l). \quad (4)$$

and scalar curvature  $\mathcal{R}$  is equal to zero.

Our goal: to compare the generalised pp-waves with purely axial torsion to the solutions of the classical models describing the interaction of gravitational and massless neutrino fields ( EW theory).

Our torsion and torsion generated curvature can be interpreted as waves traveling at speed of light.

The underlying classical pp-space of parallel Ricci curvature can then be viewed as the gravitational imprint created by a wave of some massless matter field.

We deal with the complexified curvature  $\mathfrak{R} := r(l \wedge m) \otimes (l \wedge \bar{m})$ , where  $r := \frac{1}{4}k^2 - \frac{i}{2}k'$ , hence  $R_T = \text{Re}(\mathfrak{R})$ . The curvature  $\mathfrak{R}$  is polarized, i.e.  ${}^*\mathfrak{R} = \mathfrak{R}^* = \pm i\mathfrak{R}$ , and it can be written as

$$\mathfrak{R}_{\alpha\beta\gamma\delta} = \sigma_{\alpha\beta ab} \omega^{abcd} \bar{\sigma}_{\gamma\delta cd} \quad (5)$$

Resolving (5) with respect to  $\omega$  yields

$$\omega = \xi \otimes \xi \otimes \xi \otimes \xi$$

where

$$\xi := r^{1/4} \chi \quad (6)$$

and  $\chi^a = (1, 0)$ .

The spinor (6) satisfies Weyl's (the massless Dirac) equation.

# Einstein–Weyl field equation

We consider the action as

$$S_{EW} := 2i \int \left( \xi^a \sigma^\mu{}_{ab} (\{\nabla\}_\mu \bar{\xi}^b) - (\{\nabla\}_\mu \xi^a) \sigma^\mu{}_{ab} \bar{\xi}^b \right) + K \int \mathcal{R}. \quad (7)$$

The explicit representation of the Einstein–Weyl field equations is

$$\begin{aligned} & \frac{i}{2} \left[ \sigma^\nu{}_{ab} \left( \bar{\xi}^b \{\nabla\}^\mu \xi^a - \xi^a \{\nabla\}^\mu \bar{\xi}^b \right) + \sigma^\mu{}_{ab} \left( \bar{\xi}^b \{\nabla\}^\nu \xi^a - \xi^a \{\nabla\}^\nu \bar{\xi}^b \right) \right] \\ & + i \left( \xi^a \sigma^\eta{}_{ab} (\{\nabla\}_\eta \bar{\xi}^b) g^{\mu\nu} - (\{\nabla\}_\eta \xi^a) \sigma^\eta{}_{ab} \bar{\xi}^b g^{\mu\nu} \right) - KRic^{\mu\nu} + \frac{K}{2} \mathcal{R} g^{\mu\nu} = 0, \quad (8) \end{aligned}$$

$$\sigma^\mu{}_{ab} \{\nabla\}_\mu \xi^a = 0. \quad (9)$$

The system (8), (9) has the solutions in the form of pp-waves.

# Comparison of metric–affine and Einstein–Weyl solutions

We wish to present a class of explicit solutions of EW field equations where the metric  $g$  is in the form of a pp-metric and the spinor  $\xi = r^{1/4}\xi$ .

The condition that a pp-wave needs to satisfy to be a solution of Einstein–Weyl is

$$f_{11} + f_{22} = k(x^3)^2 + \frac{2i}{K} \left( (r^{1/4})' \overline{r^{1/4}} - r^{1/4} (\overline{r^{1/4}})' \right), \quad (10)$$

The differences:

MA case: the generalised pp-wave solutions have parallel  $\{Ric\}$  curvature

EW case: the pp-wave type solutions do not necessarily have parallel Ricci curvature.

MA case: Laplacian of  $f$  can be any constant,

EW case: Laplacian of  $f$  required to be a particular constant.

The generalised pp-waves of parallel Ricci curvature are very similar to pp-type solutions of the Einstein–Weyl model.

The generalised pp-waves represent a metric-affine model for the massless neutrino.